

PHRaGMENT

Phobos Hyperspectral, Radiation, Gravity, And Magnetometry Experiment using Nanosat Technology

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Why Phobos?

- Relatively unexplored
- Limited data from past missions
- Regolith may contain ejecta
- Similarity to other small bodies
- Origin of Phobos unclear
- Potential manned mission destination, in preparation for a future manned Mars landing
- No dedicated Phobos related activities planned for Mars 2020 and 2022 missions





Motivation

"We do not know enough about the physical conditions near/at the surface of the Martian satellites low (gravity/loose regolith, etc.) ...to be able to design close proximity and surface interactions..." P-SAG Objectives (P. 24)

"...the compositions (and likely origins) of the Martian moons Phobos and Deimos may be relevant to understanding the early history of Mars" -- NRC Decadal Survey (P. 112)

"We do not have enough understanding of the geological, **compositional** and geophysical properties of Phobos and Deimos..." P-SAG Objectives (P. 24)

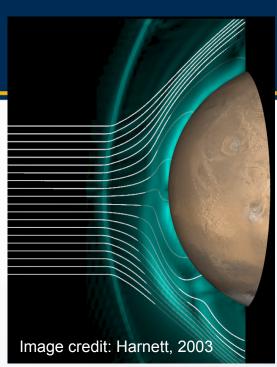
"Because Phobos and Deimos are potential staging areas and sources of resources for future human exploration of Mars, missions to Phobos and Deimos would contribute to human exploration in a way unique among missions to primitive bodies" -- NRC Decadal Survey (P. 121)

- National Research Council. *Vision and Voyages for Planetary Science in the Decade 2013-2022*. Washington, DC: The National Academies Press, 2011.
- Analysis of Strategic Knowledge Gaps Associated with Potential Human Missions to the Martian System: Final report of the Precursor Strategy Analysis Group (P-SAG), 2012



Magnetic field

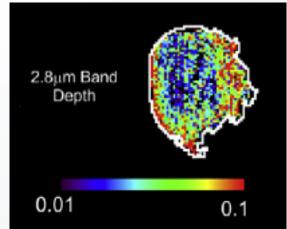
- Fobos-2 mission briefly detected faint magnetic field
- Confirmation would provide insight into the composition and origin of Phobos
- Any detected field would suggest remnant magnetization
- Characterize solar wind at Phobos
- Would observe boundary layer crossing of Mars' solar wind interaction
- Can be accomplished with (fluxgate) magnetometer and radiometer





Hyperspectral Imagery

- Provides investigation of regolith composition
- Spectral bands can be chosen to address specific theories on Phobos' origins
 - 2.3 μm band to capture clay
 - 1.9 μm band for hydrated minerals
 - 1.0 μm, 2.0 μm bands for pyroxene and olivine



Fraeman et al, Icarus, 2014

- COTS instrument options exist for NIR and visible imager
- Possible to complement with thermal, x-ray, gamma-ray, altimeter, etc.

Quasi-synchronous Orbit (QSO)

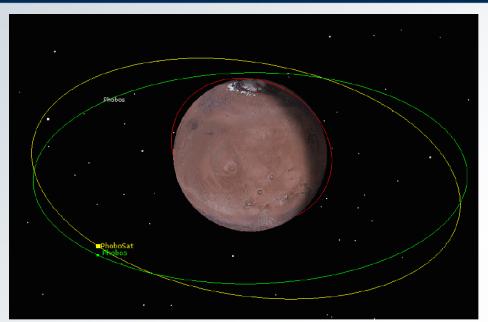


Image credit: B. Kumar, E. Srnka, S. Kuckelman, S. Chodimella, Image credit: S. Lympany, W. Jefferson, L. Smith, A. Garcia D. Ma, J. Cheung Garriga, A. Grubb, J. Paulson

- Relative distance of PHRaGMENT to Phobos varies between 100 – 400 km
- Delta-V required to reach Phobos from a representative equatorial insertion orbit > 500 m/s Georgia

Alternate orbit

- Orbit strategy also possible for a polar orbit deployment
- Revisits reduced to once every 5 orbits (23 hrs, 3 min)

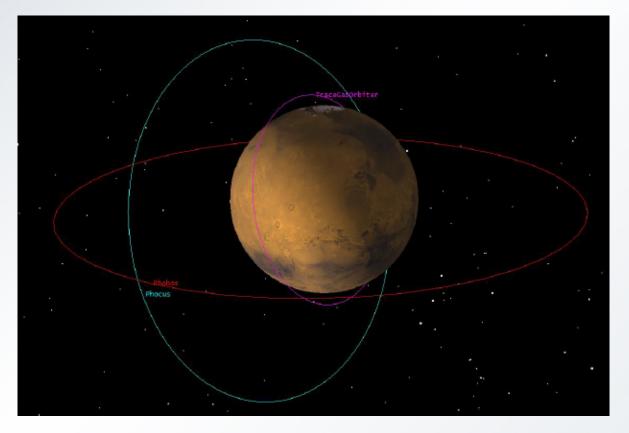


Image credit: A. Karlapudi, J. Ginn, E. Dale, L. Dressel, N. Tyman, Y. Guan, L. Lewis



Conceptual Spacecraft Design

- Both 6U and 12U configurations explored
- Comms to existing orbiting asset (e.g., MRO)

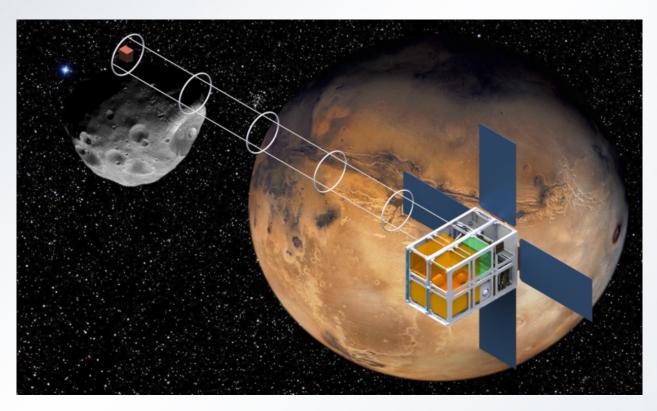


Image credit: B. Kumar, E. Srnka, S. Kuckelman, S. Chodimella, D. Ma, J. Cheung



Summary

- A 6-12U cubesat could provide significant new information on the composition and origin of Phobos
- Could provide essential data to support a future Phobos manned mission
- Use of COTS components would enable cubesat to be ready for Mars 2020 or 2022 missions
- Might provide one of the few opportunities to gather dedicated measurements of Phobos in the next decade

